

3. (Quantum Mechanics)

The purpose of this problem is to show that a spin zero particle with electric charge e in the presence of a certain radial magnetic field \mathbf{B} effectively behaves as a particle with spin $\frac{1}{2}$. The classical Lagrangian for the spin 0 particle is given by (here $\mathbf{v} = \dot{\mathbf{r}}$ and $r = |\mathbf{r}|$),

$$L = \frac{1}{2}m\mathbf{v}^2 + e\mathbf{A} \cdot \mathbf{v} \quad \mathbf{B} = \nabla \times \mathbf{A} = g \frac{\mathbf{r}}{r^3}$$

where m is the mass of the particle and g is a real parameter.

- Compute the canonical momenta \mathbf{p} conjugate to the position variables \mathbf{r} .
- Write down the Euler-Lagrange equation for the system in terms of \mathbf{r} and \mathbf{v} .
- Using the results of (b) above, show that the combination $\mathbf{L} = (L_x, L_y, L_z)$ defined by

$$\mathbf{L} = \mathbf{r} \times m\mathbf{v} - eg \frac{\mathbf{r}}{r}$$

is time-independent.

- Compute the commutators $[L_i, r_j]$ (i.e. the commutators of the components of the vectors \mathbf{L} and \mathbf{r}). An analogous result—which you are not asked to derive—for $[L_i, p_j]$ establishes that \mathbf{L} represents angular momentum.
- Compute the quantum operator L_z in spherical coordinates r, θ, ϕ using the results of (a).
- Show that the eigenvalues of L_z are half-odd-integer multiples of \hbar when the electric charge e and the parameter g are related by $eg = \frac{\hbar}{2}$.

[Hint: In a convenient gauge, the vector potential \mathbf{A} for the field \mathbf{B} is given by $\mathbf{A} = g\mathbf{n}_\phi(1 - \cos\theta)/(r \sin\theta)$ where \mathbf{n}_ϕ is the unit vector given by $\mathbf{n}_\phi = (-\sin\phi, \cos\phi, 0)$ in spherical coordinates where $x = r \sin\theta \cos\phi$, $y = r \sin\theta \sin\phi$, $z = r \cos\theta$.]

Solution:*Solution by Jonah Hyman (jthyman@g.ucla.edu)*

This problem mixes classical and quantum mechanics. It attempts to walk you step-by-step through its solution. Since we are working with a fair amount of vector algebra and calculus, it is useful to rewrite the Lagrangian in Einstein summation notation (repeated indices summed over):

$$L = \frac{1}{2}m\dot{\mathbf{r}}^2 + e\mathbf{A} \cdot \dot{\mathbf{r}} = \frac{1}{2}m\dot{r}_i\dot{r}_i + eA_i\dot{r}_i \quad (1)$$

(a) The definition of the canonical momentum is

$$\mathbf{p} = \frac{\partial L}{\partial \dot{\mathbf{r}}} \quad \text{or} \quad p_i = \frac{\partial L}{\partial \dot{r}_i} \quad (2)$$

Taking this partial derivative in (1), we get

$$\mathbf{p} = m\dot{\mathbf{r}} + e\mathbf{A} \quad \text{or} \quad p_i = m\dot{r}_i + eA_i \quad (3)$$

Here, \mathbf{A} is the vector potential for the field \mathbf{B} given in the hint. In other words,

$$\boxed{\mathbf{p} = m\mathbf{v} + e\mathbf{A}} \quad (4)$$

(b) The Euler-Lagrange equation for this system is just the Newton's second law equation for the system. This is just a particle of charge e in a magnetic field, so the Newton's second law equation is given by the Lorentz force law:

$$m\mathbf{a} = e\mathbf{v} \times \mathbf{B} \quad (5)$$

If you (like the author) didn't think of this, you can derive this from the definition of the Euler-Lagrange equation:

$$\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{\mathbf{r}}} \right) - \frac{\partial L}{\partial \mathbf{r}} = 0 \quad \text{or} \quad \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{r}_i} \right) - \frac{\partial L}{\partial r_i} = 0 \quad (6)$$

The first term is equal to $\frac{d\mathbf{p}}{dt}$. Using (3) and taking a *total* time derivative (applying the chain rule to the vector potential $\mathbf{A}(\mathbf{r})$), we get

$$\begin{aligned} \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{r}_i} \right) &= \frac{d}{dt} (m\dot{r}_i + eA_i) \\ &= m\ddot{r}_i + e (\partial_j A_i) \dot{r}_j \end{aligned} \quad (7)$$

The other term in the Euler-Lagrange equation is given by

$$\frac{\partial L}{\partial r_i} = e (\partial_i A_j) \dot{r}_j \quad (8)$$

Thus, the Euler-Lagrange equation is

$$\begin{aligned} 0 &= m\ddot{r}_i + e (\partial_j A_i) \dot{r}_j - e (\partial_i A_j) \dot{r}_j \\ \implies m\ddot{r}_i &= e [(\partial_i A_j) \dot{r}_j - (\partial_j A_i) \dot{r}_j] \end{aligned} \quad (9)$$

The quantity in brackets might remind you of the right-hand side of the BAC-CAB identity:

$$\mathbf{a} \times (\mathbf{b} \times \mathbf{c}) = \mathbf{b}(\mathbf{a} \cdot \mathbf{c}) - \mathbf{c}(\mathbf{a} \cdot \mathbf{b}) \quad \text{or} \quad (\mathbf{a} \times (\mathbf{b} \times \mathbf{c}))_i = b_i a_j c_j - c_i a_j b_j$$

Pattern-matching to (9), and noting that $\partial_i \dot{r}_j = 0$, we can identify $a_i \longleftrightarrow \dot{r}_i$, $b_i \longleftrightarrow \partial_i$, and $c_i \longleftrightarrow A_i$. This means that (9) becomes

$$m\ddot{r}_i = e [\dot{\mathbf{r}}_i \times (\nabla \times \mathbf{A})]_i \quad (10)$$

Since $\ddot{\mathbf{r}} = \mathbf{a}$ (acceleration), $\dot{\mathbf{r}} = \mathbf{v}$ (velocity), and $\nabla \times \mathbf{A} = \mathbf{B}$, this becomes

$$m\mathbf{a} = e\mathbf{v} \times \mathbf{B} \quad (11)$$

In this problem, $\mathbf{B} = g\frac{\mathbf{r}}{r^3}$. In terms of \mathbf{r} and \mathbf{b} , the Euler-Lagrange equation boils down to

$$m\dot{\mathbf{v}} = eg\mathbf{v} \times \frac{\mathbf{r}}{r^3} \quad (12)$$

- (c) To show that \mathbf{L} is time-independent, we need only show that its *total* time derivative is zero. Using the product and chain rules, we get

$$\begin{aligned} \frac{d\mathbf{L}}{dt} &= \frac{d}{dt} \left(\mathbf{r} \times m\mathbf{v} - eg\frac{\mathbf{r}}{r} \right) \\ &= \frac{d\mathbf{r}}{dt} \times m\mathbf{v} + \mathbf{r} \times m\frac{d\mathbf{v}}{dt} - eg\frac{1}{r}\frac{d\mathbf{r}}{dt} - eg\mathbf{r}\frac{d}{dt}\left(\frac{1}{r}\right) \\ &= \mathbf{v} \times m\mathbf{v} + \mathbf{r} \times m\dot{\mathbf{v}} - eg\frac{\mathbf{v}}{r} + eg\frac{\mathbf{r}}{r^2}\frac{dr}{dt} \\ \frac{d\mathbf{L}}{dt} &= \mathbf{r} \times m\dot{\mathbf{v}} - eg\frac{\mathbf{v}}{r} + eg\frac{\mathbf{r}}{r^2}\frac{dr}{dt} \quad \text{since } \mathbf{v} \times \mathbf{v} = 0 \end{aligned} \quad (13)$$

The problem asks us to use our answer from part (b). Using (12) and substituting for $m\dot{\mathbf{v}}$, we get

$$\begin{aligned} \frac{d\mathbf{L}}{dt} &= \mathbf{r} \times \left(eg\mathbf{v} \times \frac{\mathbf{r}}{r^3} \right) - eg\frac{\mathbf{v}}{r} + eg\frac{\mathbf{r}}{r^2}\frac{dr}{dt} \\ &= eg \left[\frac{1}{r^3}\mathbf{r} \times (\mathbf{v} \times \mathbf{r}) - \frac{\mathbf{v}}{r} + \frac{\mathbf{r}}{r^2}\frac{dr}{dt} \right] \end{aligned} \quad (14)$$

The triple product $\mathbf{r} \times (\mathbf{v} \times \mathbf{r})$ can be simplified using the BAC-CAB rule:

$$\mathbf{a} \times (\mathbf{b} \times \mathbf{c}) = \mathbf{b}(\mathbf{a} \cdot \mathbf{c}) - \mathbf{c}(\mathbf{a} \cdot \mathbf{b})$$

This gets us

$$\begin{aligned} \mathbf{r} \times (\mathbf{v} \times \mathbf{r}) &= \mathbf{v}(\mathbf{r} \cdot \mathbf{r}) - \mathbf{r}(\mathbf{r} \cdot \mathbf{v}) \\ &= r^2\mathbf{v} - \mathbf{r}(\mathbf{r} \cdot \mathbf{v}) \end{aligned} \quad (15)$$

Plugging this back into (14), we get

$$\begin{aligned} \frac{d\mathbf{L}}{dt} &= eg \left[\frac{1}{r^3} (r^2\mathbf{v} - \mathbf{r}(\mathbf{r} \cdot \mathbf{v})) - \frac{\mathbf{v}}{r} + \frac{\mathbf{r}}{r^2}\frac{dr}{dt} \right] \\ &= eg \left[\frac{\mathbf{v}}{r} - \frac{\mathbf{r}}{r^3}(\mathbf{r} \cdot \mathbf{v}) - \frac{\mathbf{v}}{r} + \frac{\mathbf{r}}{r^2}\frac{dr}{dt} \right] \\ &= eg \left[-\frac{\mathbf{r}}{r^3}(\mathbf{r} \cdot \mathbf{v}) + \frac{\mathbf{r}}{r^2}\frac{dr}{dt} \right] \end{aligned} \quad (16)$$

One way to show that the term in brackets vanishes is to remember that the radial component of the velocity vector is the change in the radius:

$$\hat{\mathbf{r}} \cdot \mathbf{v} = \frac{dr}{dt} \quad (17)$$

Then,

$$\frac{\mathbf{r}}{r^2}\frac{dr}{dt} = \frac{\mathbf{r}}{r^2}(\hat{\mathbf{r}} \cdot \mathbf{v}) = \frac{\mathbf{r}}{r^3}(\mathbf{r} \cdot \mathbf{v}) \quad (18)$$

Another way to show that the term in brackets vanishes is to take the antiderivative of $\mathbf{r} \cdot \mathbf{v}$:

$$\begin{aligned}
 \frac{\mathbf{r}}{r^3} (\mathbf{r} \cdot \mathbf{v}) &= \frac{\mathbf{r}}{r^3} \left(\mathbf{r} \cdot \frac{d\mathbf{r}}{dt} \right) \\
 &= \frac{\mathbf{r}}{r^3} \left(\frac{1}{2} \frac{d}{dt} (\mathbf{r} \cdot \mathbf{r}) \right) \\
 &= \frac{\mathbf{r}}{r^3} \left(\frac{1}{2} \frac{d}{dt} (r^2) \right) \\
 &= \frac{\mathbf{r}}{r^3} \left(r \frac{dr}{dt} \right) \\
 &= \frac{\mathbf{r}}{r^2} \frac{dr}{dt}
 \end{aligned} \tag{19}$$

Either way, we get that $\frac{\mathbf{r}}{r^3} (\mathbf{r} \cdot \mathbf{v}) = \frac{\mathbf{r}}{r^2} \frac{dr}{dt}$, so $-\frac{\mathbf{r}}{r^3} (\mathbf{r} \cdot \mathbf{v}) + \frac{\mathbf{r}}{r^2} \frac{dr}{dt} = 0$. Plugging this result into (16), we establish the result

$$\frac{d\mathbf{L}}{dt} = 0 \tag{20}$$

so \mathbf{L} is time-independent.

(d) To calculate the commutators $[L_i, r_j]$, we need to use the canonical commutation relations

$$[r_i, r_j] = 0 \quad \text{and} \quad [r_i, p_j] = i\hbar \delta_{ij} \tag{21}$$

This means that we need to write \mathbf{L} in terms of \mathbf{r} and \mathbf{p} . Since all components of \mathbf{r} commute with each other, any terms in \mathbf{L} that are independent of \mathbf{p} commute with \mathbf{r} automatically. From part (a), $\mathbf{p} = m\mathbf{v} + e\mathbf{A}$, and \mathbf{A} depends only on \mathbf{r} , so

$$\begin{aligned}
 \mathbf{L} &= \mathbf{r} \times m\mathbf{v} - eg \frac{\mathbf{r}}{r} \\
 &= \mathbf{r} \times (\mathbf{p} - e\mathbf{A}) - eg \frac{\mathbf{r}}{r} \\
 &= \mathbf{r} \times \mathbf{p} - e \mathbf{r} \times \mathbf{A} - eg \frac{\mathbf{r}}{r}
 \end{aligned} \tag{22}$$

In components, this becomes

$$L_i = \epsilon_{ijk} r_j p_k + f_i(\mathbf{r}) \tag{23}$$

where ϵ_{ijk} is the Levi-Civita symbol and $f_i(\mathbf{r})$ is a function of \mathbf{r} . Since f_i is only a function of \mathbf{r} , this term commutes with all components of \mathbf{r} . Therefore, applying the commutation rules and relations, we get

$$\begin{aligned}
 [L_i, r_j] &= [\epsilon_{ik\ell} r_k p_\ell, r_j] + [f_i(\mathbf{r}), r_j] \\
 &= [\epsilon_{ik\ell} r_k p_\ell, r_j] \quad \text{since } [r_i, r_j] = 0 \\
 &= \epsilon_{ik\ell} r_k [p_\ell, r_j] + \epsilon_{ik\ell} [r_k, r_j] p_\ell \quad \text{as } [AB, C] = A[B, C] + [A, C]B \\
 &= \epsilon_{ik\ell} r_k [p_\ell, r_j] \quad \text{since } [r_i, r_j] = 0 \\
 &= -\epsilon_{ik\ell} r_k [r_j, p_\ell] \quad \text{since } [A, B] = -[B, A] \\
 &= -\epsilon_{ik\ell} r_k (i\hbar \delta_{j\ell}) \quad \text{since } [r_i, p_j] = i\hbar \delta_{ij} \\
 &= -i\hbar \epsilon_{ikj} r_k \\
 &= i\hbar \epsilon_{ijk} r_k \quad \text{since } \epsilon_{ijk} = -\epsilon_{ikj}
 \end{aligned} \tag{24}$$

We have found that

$$\boxed{[L_i, r_j] = i\hbar \epsilon_{ijk} r_k} \tag{25}$$

- (e) To write the quantum operators in spherical coordinates, we need to use the canonical quantization prescription

$$p_x = \frac{\hbar}{i} \frac{\partial}{\partial x} \quad \text{and} \quad p_y = \frac{\hbar}{i} \frac{\partial}{\partial y} \quad (26)$$

From (22), we can write L_z in terms of x , y , p_x , and p_y :

$$\begin{aligned} L_z &= (\mathbf{r} \times \mathbf{p})_z - e (\mathbf{r} \times \mathbf{A})_z - eg \left(\frac{\mathbf{r}}{r} \right)_z \\ &= (xp_y - yp_x) - e (xA_y - yA_x) - eg \frac{z}{r} \end{aligned} \quad (27)$$

Canonically quantizing the first term, we get

$$xp_y - yp_x = \frac{\hbar}{i} \left(x \frac{\partial}{\partial y} - y \frac{\partial}{\partial x} \right) \quad (28)$$

The best way to convert this to spherical coordinates is to guess (from previous experience) that the partial derivative $\frac{\partial}{\partial \varphi}$ is involved. In spherical coordinates, we have

$$x = r \sin \theta \cos \varphi \quad \text{and} \quad y = r \sin \theta \sin \varphi \quad (29)$$

Thus,

$$\begin{aligned} \frac{\partial}{\partial \varphi} &= \frac{\partial x}{\partial \varphi} \frac{\partial}{\partial x} + \frac{\partial y}{\partial \varphi} \frac{\partial}{\partial y} \\ &= -r \sin \theta \sin \varphi \frac{\partial}{\partial x} + r \sin \theta \cos \varphi \frac{\partial}{\partial y} \\ &= -y \frac{\partial}{\partial x} + x \frac{\partial}{\partial y} \end{aligned} \quad (30)$$

Therefore, plugging into (28), we get

$$xp_y - yp_x = \frac{\hbar}{i} \left(\frac{\partial}{\partial \varphi} \right) = -i\hbar \frac{\partial}{\partial \varphi} \quad (31)$$

For the second term in (27), we need to use the vector potential given in the hint:

$$\begin{aligned} \mathbf{A} &= g\hat{\varphi} \frac{1 - \cos \theta}{r \sin \theta} \\ &= g (-\sin \varphi \hat{\mathbf{x}} + \cos \varphi \hat{\mathbf{y}}) \frac{1 - \cos \theta}{r \sin \theta} \end{aligned} \quad (32)$$

Therefore, in spherical coordinates, the second term in (27) simplifies to

$$\begin{aligned} -e (xA_y - yA_x) &= -e \left[(r \sin \theta \cos \varphi) \left(g \cos \varphi \frac{1 - \cos \theta}{r \sin \theta} \right) - (r \sin \theta \sin \varphi) \left(-g \sin \varphi \frac{1 - \cos \theta}{r \sin \theta} \right) \right] \\ &= -eg r \sin \theta \frac{1 - \cos \theta}{r \sin \theta} [\cos^2 \varphi + \sin^2 \varphi] \\ &= -eg(1 - \cos \theta) \end{aligned} \quad (33)$$

The last term in (27) simplifies to

$$\begin{aligned} -eg \frac{z}{r} &= -eg \frac{r \cos \theta}{r} \\ &= -eg \cos \theta \end{aligned} \quad (34)$$

Putting all three terms together using (31), (33), and (34), we get

$$L_z = -i\hbar \frac{\partial}{\partial \varphi} - eg(1 - \cos \theta) - eg \cos \theta$$

$$\boxed{L_z = -i\hbar \frac{\partial}{\partial \varphi} - eg} \quad (35)$$

(f) If $eg = \frac{\hbar}{2}$, then

$$L_z = \hbar \left(-i \frac{\partial}{\partial \varphi} - \frac{1}{2} \right) \quad (36)$$

To find the eigenvalues of this operator, note that any wave function must be single-valued upon taking $\varphi \rightarrow \varphi + 2\pi$. Therefore, by Fourier analysis, any wave function must be a linear combination of functions of the form $e^{in\varphi}$, where n is an integer. Applying the operator L_z to such a function, we get

$$\begin{aligned} L_z e^{in\varphi} &= \hbar \left(-i \frac{\partial}{\partial \varphi} - \frac{1}{2} \right) e^{in\varphi} \\ &= \hbar \left(-i(in) - \frac{1}{2} \right) e^{in\varphi} \\ &= \hbar \left(n - \frac{1}{2} \right) e^{in\varphi} \\ &= \hbar \underbrace{\left(\frac{2n-1}{2} \right)}_{\text{half odd-integer}} e^{in\varphi} \end{aligned} \quad (37)$$

This means that $e^{in\varphi}$ is an eigenfunction of L_z with eigenvalue $\hbar \left(\frac{2n-1}{2} \right)$. Thus, when $eg = \frac{\hbar}{2}$, the eigenvalues of L_z are half-odd-integer multiples of \hbar .

This result implies that in the presence of this magnetic field, this spin-zero particle acts like a spin- $\frac{1}{2}$ particle, in that its L_z angular momentum eigenvalues are half-odd-integer multiples of \hbar .