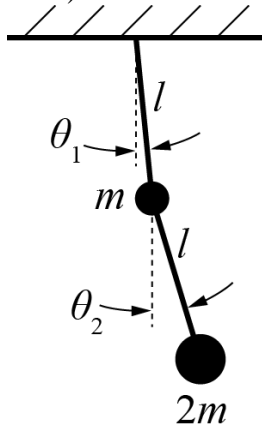


Physics & Astronomy Comprehensive Exam, UCLA, Fall 2013*1. Classical Mechanics*

Consider the double pendulum shown below with equal string lengths l and unequal masses m and $2m$, (constrained to move in the plane shown, gravitational acceleration g). Find the frequencies of the normal modes for small oscillations about the equilibrium position (θ_1 and $\theta_2 \ll 1$).



*Questions for the Comprehensive Exam Fall 2013**2. Classical Mechanics*

A photon of energy E_1 collides at an angle θ with another photon of energy E_2 . Find the minimum value of E_1 (given E_2 and θ) permitting the formation of a pair of particles of mass m . With this expression, calculate the minimum energy a gamma ray must have to create an electron-positron pair by colliding with a typical photon from the cosmic microwave background (one significant figure is sufficient). [Treat the photon here as a relativistic quantum particle.]

*Questions for the Comprehensive Exam Fall 2013**3. Quantum Mechanics: Scattering*

Consider the scattering of a spinless particle of mass m from a diatomic molecule. The incoming particle travels along the z -axis. Assume that the molecule is much heavier than the scattering particle and that there is no recoil. The two atoms in the molecule are aligned along the y -axis and localized at $y = +b$ and $y = -b$. The potential the particle feels in the presence of the molecule can be modeled by the following potential:

$$V(\vec{x}) = \alpha(\delta(y-b)\delta(x)\delta(z) + \delta(y+b)\delta(x)\delta(z))$$

- (a) Calculate the scattering amplitude in the first Born approximation.
- (b) Calculate the differential cross section from (a) (Express the result in terms of the scattering angles).
- (c) Calculate the total cross section. Do the integrals exactly. You might find the following integrals helpful:

$$\int_0^{2\pi} d\alpha |\cos(x \sin \alpha)|^2 = \pi(1 + J_0(2x)), \quad \int_0^\pi d\alpha (\sin \alpha) J_0(x \sin \alpha) = \frac{a \sin x}{x}$$

Name:

Questions for the Comprehensive Exam Fall 2013

4. Quantum Mechanics

Consider a system of two spin 1/2 particles, labeled a and b , with respective spin operators \mathbf{S}_a and \mathbf{S}_b . We ignore all quantum numbers but those of spin. The particles are in the state $|\Psi\rangle$ of zero total angular momentum, which we consider normalized to $\langle\Psi|\Psi\rangle = 1$. Let \mathbf{n}_a and \mathbf{n}_b be two independent unit vectors. Compute the expectation value of the product of the spin operators projected onto the directions \mathbf{n}_a and \mathbf{n}_b respectively, namely, $\langle\Psi|(\mathbf{n}_a \cdot \mathbf{S}_a)(\mathbf{n}_b \cdot \mathbf{S}_b)|\Psi\rangle$.

*Questions for the Comprehensive Exam Fall 2013**5. Quantum Mechanics*

Consider a one dimensional simple harmonic oscillator with the usual Hamiltonian,

$$H_0 = \frac{p^2}{2m} + \frac{1}{2}m\omega^2 x^2$$

Now consider the effect of a perturbation,

$$V = \frac{1}{2}\varepsilon m\omega^2 x^2 \quad \text{with } \varepsilon \ll 1.$$

Using perturbation theory, find the new ground state key $|0\rangle$ to order ε and the ground state energy shift to order ε^2 . Solve this problem exactly and compare with the results obtained using perturbation theory. You may assume without proof that

$$\langle u_{n'} | x | u_n \rangle = \sqrt{\frac{\hbar}{2m\omega}} (\sqrt{n+1}\delta_{n',n+1} + \sqrt{n}\delta_{n',n-1})$$

where $|u_n\rangle$ is the n^{th} eigenstate of H_0 .

Questions for the Comprehensive Exam Fall 2013

6. *Quantum Mechanics: Perturbation theory time (in)dependent*

The Hamiltonian of a particle of mass m in a 1D finite well is given by

$$H_0 = \frac{p^2}{2m} + V(x), \quad V(x) = \begin{cases} 0 & 0 < x < L \\ \infty & \text{otherwise} \end{cases}$$

A time dependent perturbation is added

$$H_{total} = H_0 + H', \quad H' = \lambda \delta(x - L/2) f(t)$$

where λ is a constant and $f(t)$ is a time dependent function.

- (a) Calculate the matrix elements of H' with the eigenstates of the unperturbed Hamiltonian.
- (b) In this part of the problem we consider a time independent perturbation $f(t) = 1$. First, calculate the first nonzero correction of the ground state energy. Second, at which order (if any) will the first excited state receive a nonzero correction? Please back up your answer with an argument/calculation.
- (c) Now take the following time dependent function

$$f(t) = \int_{-\infty}^{\infty} d\omega \rho(\omega) (e^{i\omega t} + e^{-i\omega t}) \quad \text{with} \quad \rho(\omega) = \sqrt{\frac{\alpha}{\pi}} e^{-\alpha\omega^2}$$

If the system is in its ground state at time $t = 0$, what are the possible transitions into excited states that the system can make at time $t > 0$? Use first order time-dependent perturbation theory.

- (d) Find the transition rate into these excited states using Fermi's golden rule.

*Questions for the Comprehensive Exam Fall 2013**7. Quantum Mechanics: Jaynes-Cummings Model*

A spin-1/2 particle (bare eigenstates $|\uparrow\rangle$ and $|\downarrow\rangle$) is constrained to move in a 1D harmonic potential (bare eigenstates $|n\rangle$ for $n \in \{0,1,2,\dots\}$). In the absence of coupling, the Hamiltonian is given by

$$H_0 = \hbar\omega_0(\hat{a}^\dagger\hat{a} + 1/2) + \hbar\omega_q\hat{\sigma}_z/2$$

where $\hbar\omega_0$ is the harmonic oscillation energy spacing and $\hbar\omega_q$ is the energy difference between $|\uparrow\rangle$ and $|\downarrow\rangle$ in the absence of coupling. Consider the case where the particle's spin is coupled to its motion through the interaction Hamiltonian

$$H_{\text{int}} = \hbar\Omega(\hat{a}^\dagger\hat{\sigma}_- + \hat{a}\hat{\sigma}_+)/2$$

where Ω is the coupling strength and the operators for the spin degree of freedom are defined by $\hat{\sigma}_z = |\uparrow\rangle\langle\uparrow| - |\downarrow\rangle\langle\downarrow|$, $\hat{\sigma}_+ = |\uparrow\rangle\langle\downarrow|$, and $\hat{\sigma}_- = |\downarrow\rangle\langle\uparrow|$.

(a) Let $\Delta \equiv \omega_0 - \omega_q > 0$. Consider the “dispersive regime” defined by $\Delta \gg \Omega/2$. Find the energies and eigenstates of $H = H_0 + H_{\text{int}}$. You may treat H_{int} as a perturbation and keep only the first nonzero-order terms, and you may consider coupling between nearly-degenerate states only. Do not bother to normalize the eigenstates.

(b) Consider the “resonant regime” where $\Delta \rightarrow 0$ and ω_0 and ω_q are both equal to some frequency ω . Find the energies and eigenstates of $H = H_0 + H_{\text{int}}$. You may treat H_{int} as a perturbation and use first-order degenerate perturbation theory, considering coupling between nominally degenerate states only. Hint: consider the state $|\psi\rangle = \frac{1}{\sqrt{2}}(|\downarrow, n\rangle + e^{i\phi}|\uparrow, n-1\rangle)$.

*Questions for the Comprehensive Exam Fall 2013**8. Statistical Mechanics*

We propose to evaluate the **Richardson** effect, namely the electric current density of electrons which is produced by heating up a metal in the presence of an external electric potential. The potential energy of an electron just outside the metal is denoted $W > 0$. The potential energy for electrons inside the metal is taken to be 0. The electrons are considered otherwise non-interacting, and filled up to chemical potential μ with $\mu < W$. Since we consider the problem to be at sufficiently low temperature, μ may be identified with the Fermi energy.

- (a) State the condition on the momentum of an electron that can escape from the metal to the outside as a function of W and μ .
- (b) Derive a general expression for the current density I of electrons leaving the metal.
- (c) Obtain an approximation of your result in (b) valid for sufficiently low temperatures.

*Questions for the Comprehensive Exam Fall 2013**9. Statistical Mechanics*

The flexing modes of a thin plate at wave number k have an angular frequency ω such that:

$\omega^2 = \gamma k^4$. Consider such waves propagating in one dimension around a thin ring of radius R .

What is the contribution to the heat capacity of these azimuthal modes? The plate is in thermal equilibrium at a low temperature T . You may write your answer in terms of:

$$Z(x) = \int_{-\infty}^{+\infty} \frac{y^x}{e^y - 1} dy \quad \text{where } x \text{ is a pure number.}$$

Name:

Questions for the Comprehensive Exam Fall 2013

10. Statistical Mechanics

A liquid is in equilibrium with its vapor at temperature and pressure: T_v ; P_v . The surface between liquid and vapor is flat. The temperature of the vapor is increased to $T_v + \Delta T$ while keeping its pressure fixed. The liquid remains at temperature and pressure: T_v ; P_v . Evaluate the net flux of gas to the liquid. You may treat the vapor as an ideal noble gas with atoms of mass m .

*Questions for the Comprehensive Exam Fall 2013**11. Electromagnetism: A spherical capacitor*

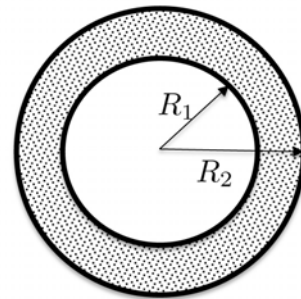
The gap between two spherical conducting shells of radii $R_1 < R_2$ is filled with a spatially inhomogeneous dielectric. As a result, the dielectric constant depends on the polar angle θ (measured from the north pole of the sphere) as:

$$\varepsilon(\theta) = \varepsilon_1 + \varepsilon_2 \cos^4 \theta$$

Both ε_1 and ε_2 are constants.

(a) When charged so that the inner and outer spheres have charges $+Q$ and $-Q$ respectively, show that the internal electric field is purely radial. Determine how that radial electric field E_r depends on the polar angles, θ and ϕ .

(b) Calculate the capacitance C of the system.



*Questions for the Comprehensive Exam Fall 2013**12. Electromagnetism*

Assume the existence of magnetic charge related to the magnetic field by the local relation

$$\vec{\nabla} \cdot \vec{B} = \mu_0 \rho_m.$$

- (a). Using the Gauss's theorem, obtain the magnetic field \vec{B} of a point magnetic charge at the origin.
- (b). In the absence of the magnetic charge, the curl of the electric field is given by the Faraday's law, $\vec{\nabla} \times \vec{E} = -\frac{\partial \vec{B}}{\partial t}$. Show that this law is incompatible with the magnetic charge density that is a function of time.
- (c). Assuming that magnetic charge is conserved, derive the local relation between the magnetic charge current density \vec{J}_m and the magnetic density ρ_m .
- (d). Modify Faraday's law as given in part (b) to obtain a law consistent with the presence of the magnetic charge density that is a function of position and time.

Name:

Questions for the Comprehensive Exam Fall 2013

13. Electromagnetism

An un-polarized plane electromagnetic wave is scattered by a free electron. Derive the differential cross-section for scattering in the non-relativistic limit (Thompson scattering).

Name:

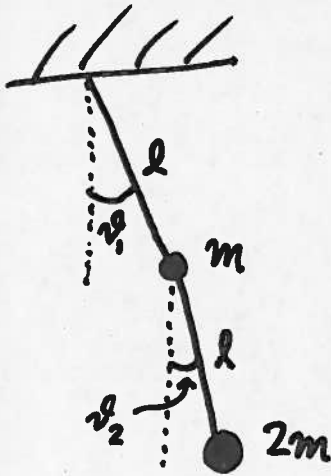
Questions for the Comprehensive Exam Fall 2013

14. Electromagnetism

A plane electromagnetic wave of frequency ω is normally incident on an infinitesimally-thin planar sheet with a resistivity $Z_s = Z_0 = 377$ ohms per square ($4\pi/c$ in cgs units). The medium on both sides of the sheet is vacuum, and the sheet is non-magnetic.

- (a) Calculate the power coefficients of reflection (R), absorption (A), and transmission (T).
- (b) What value of the resistivity Z_m would maximize A ? What is the maximal value of A , and why is this the limit?

Classical Mechanics Q1 : double pendulum



Consider the double pendulum with unequal masses shown at left. For small oscillations about the equilibrium position $\theta_1 \rightarrow 0$, $\theta_2 \rightarrow 0$, find the normal mode frequencies. You may assume $\theta_1 \ll 1$, $\theta_2 \ll 1$.

$$K = \frac{1}{2} m \left[l \frac{d}{dt} \sin(\theta_1) \right]^2 + \frac{1}{2} \cdot 2m \left[l \frac{d}{dt} \sin(\theta_1) + l \frac{d}{dt} \sin(\theta_2) \right]^2 \quad (1)$$

$$\approx \frac{1}{2} m l^2 \dot{\theta}_1^2 + \frac{1}{2} m l^2 [2 \dot{\theta}_1^2 + 4 \dot{\theta}_1 \dot{\theta}_2 + 2 \dot{\theta}_2^2]$$

$$= \frac{1}{2} m l^2 (3 \dot{\theta}_1^2 + 4 \dot{\theta}_1 \dot{\theta}_2 + 2 \dot{\theta}_2^2)$$

$$V = (m + 2m)g(l - l \cos(\theta_1)) + 2mg(l - l \cos(\theta_2))$$

$$\approx \frac{3}{2} mgl \theta_1^2 + mgl \theta_2^2$$

$$\therefore L = \frac{1}{2} m l^2 (3 \dot{\theta}_1^2 + 4 \dot{\theta}_1 \dot{\theta}_2 + 2 \dot{\theta}_2^2) - \frac{3}{2} mgl \theta_1^2 - mgl \theta_2^2$$

The momenta are

$$p_1 = \frac{\partial L}{\partial \dot{\theta}_1} = 3ml^2 \dot{\theta}_1 + 2ml^2 \dot{\theta}_2$$

$$p_2 = \frac{\partial L}{\partial \dot{\theta}_2} = 2ml^2 \dot{\theta}_2 + 2ml^2 \dot{\theta}_1$$

we can write the equations of motion:

$$\frac{dp_1}{dt} = \frac{\partial L}{\partial \theta_1}$$

$$\frac{d}{dt} [3ml^2 \dot{\theta}_1 + 2ml^2 \dot{\theta}_2] = -3mg l \theta_1$$

$$(I) \quad 3\ddot{\theta}_1 + 2\ddot{\theta}_2 = -3\frac{g}{l} \theta_1$$

$$\text{and} \quad \frac{dp_2}{dt} = \frac{\partial L}{\partial \theta_2}$$

(2)

$$\frac{d}{dt} [2\eta l^2 \dot{x}_2 + 2\eta l^2 \dot{x}_1] = -2\eta g l x_2$$

$$2\ddot{x}_2 + 2\ddot{x}_1 = -2\frac{g}{l} x_2$$

$$(II) \quad \ddot{x}_2 + \ddot{x}_1 = -\frac{g}{l} x_2$$

assume $x_1(t) = A_1 e^{i\omega t}$ $x_2(t) = A_2 e^{i\omega t}$

from (I) we have

$$-\omega^2 A_1 \cdot 3 - 2\omega^2 A_2 = -3A_1 \frac{g}{l}$$

$$A_1 \left(3\frac{g}{l} - 3\omega^2 \right) = 2\omega^2 A_2$$

$$\therefore A_2 = A_1 \left(\frac{3}{2} \frac{g}{l} \frac{1}{\omega^2} - \frac{3}{2} \right)$$

from (II) we have

$$-A_2 \omega^2 - A_1 \omega^2 = -A_2 \frac{g}{l}$$

$$A_1 \omega^2 = A_2 \left(\frac{g}{l} - \omega^2 \right)$$

$$= A_1 \left(\frac{3}{2} \frac{g}{l} \frac{1}{\omega^2} - \frac{3}{2} \right) \left(\frac{g}{l} - \omega^2 \right)$$

$$\therefore \omega^2 = \frac{3}{2} \left(\frac{g}{l} \right)^2 \frac{1}{\omega^2} - \underbrace{\frac{3}{2} \frac{g}{l} - \frac{3}{2} \frac{g}{l}}_{-3\frac{g}{l}} + \frac{3}{2} \omega^2$$

$$0 = \frac{3}{2} \left(\frac{g}{l} \right)^2 \frac{1}{\omega^2} - 3\frac{g}{l} + \frac{1}{2} \omega^2$$

$$0 = \frac{3}{2} \left(\frac{g}{l} \right)^2 - 3\frac{g}{l} \omega^2 + \frac{1}{2} \omega^4$$

$$0 = \frac{3}{2} - 3\frac{l}{g} \omega^2 + \frac{1}{2} \left(\frac{l}{g} \omega^2 \right)^2$$

③

$$\text{So } \frac{1}{g}\omega^2 = \frac{3 \pm \sqrt{9-3}}{1} = 3 \pm \sqrt{6}$$

$$\omega_+ = \sqrt{\frac{g}{l}} \sqrt{3 + \sqrt{6}}$$

$$\omega_- = \sqrt{\frac{g}{l}} \sqrt{3 - \sqrt{6}}$$

Solution 2

Matrix method

(4)

$$K = \frac{1}{2} m l^2 (3\ddot{\theta}_1^2 + 4\ddot{\theta}_1\ddot{\theta}_2 + 2\ddot{\theta}_2^2)$$

$$V = \frac{3}{2} m g l \theta_1^2 + m g l \theta_2^2$$
$$= \frac{1}{2} \cdot 3 m g l \cdot \theta_1^2 + \frac{1}{2} \cdot 2 m g l \cdot \theta_2^2$$

we identify

$$M_{11} = 3 m l^2$$

$$M_{22} = 2 m l^2$$

$$M_{12} = M_{21} = 2 m l^2$$

$$\tilde{M} = \begin{pmatrix} 3 m l^2 & 2 m l^2 \\ 2 m l^2 & 2 m l^2 \end{pmatrix} = m l^2 \begin{pmatrix} 3 & 2 \\ 2 & 2 \end{pmatrix}$$

$$K_{11} = 3 m g l$$

$$K_{22} = 2 m g l$$

$$K_{12} = K_{21} = 0$$

$$\tilde{K} = m g l \begin{pmatrix} 3 & 0 \\ 0 & 2 \end{pmatrix}$$

we are looking for the solution to

$$(\tilde{K} - \lambda \tilde{M}) \vec{a} = 0 \quad \lambda \equiv \omega^2$$

we construct the characteristic equation by taking the determinant:

$$\begin{vmatrix} 3 m g l - \omega^2 \cdot 3 m l^2 & -\omega^2 2 m l^2 \\ -\omega^2 2 m l^2 & 2 m g l - \omega^2 \cdot 2 m l^2 \end{vmatrix} = 0$$

$$(3g - 3l\omega^2)(2g - 2l\omega^2) - 4l^2\omega^4 = 0$$

$$6g^2 - 6gl\omega^2 - 6gl\omega^2 + 6l^2\omega^4 - 4l^2\omega^4 = 0$$

⑤

$$2l^2\omega^4 - 12gl\omega^2 + 6g^2 = 0$$

$$\frac{l^2}{g^2}\omega^4 - 6\frac{l}{g}\omega^2 + 3 = 0$$

$$\text{so } \omega^2 \frac{l}{g} = \frac{6 \pm \sqrt{36 - 12}}{2} = 3 \pm \sqrt{6}$$

$$\omega_+ = \sqrt{\frac{g}{l}} \sqrt{3 + \sqrt{6}}$$

$$\omega_- = \sqrt{\frac{g}{l}} \sqrt{3 - \sqrt{6}}$$

Q2

Without loss of generality we take photon 1 to be traveling in the \hat{z} direction and photon 2 to have its momentum in the XZ plane. Then the 4-momenta are

$$p_1 = (p_x, p_y, p_z, \frac{E}{c}) = (0, 0, \frac{E_1}{c}, \frac{E_1}{c})$$

$$p_2 = (\frac{\sin\theta E_2}{c}, 0, \frac{\cos\theta E_2}{c}, \frac{E_2}{c})$$

$$p_t = p_1 + p_2 = (\frac{\sin\theta E_2}{c}, 0, \frac{\cos\theta E_2 + E_1}{c}, \frac{E_1 + E_2}{c})$$

$$\begin{aligned} \text{The total } |\vec{p}| &= \frac{1}{c} \sqrt{\sin^2\theta E_2^2 + \cos^2\theta E_2^2 + E_1^2 + 2E_1 E_2 \cos\theta} \\ &= \frac{1}{c} \sqrt{E_1^2 + E_2^2 + 2E_1 E_2 \cos\theta} \end{aligned}$$

At threshold $m_3 = m_4$ are not moving relative to each other. \Rightarrow They have the same momenta.

$$p_3 = p_4 \text{ such that } |\vec{p}| = p \text{ and } E = \sqrt{p^2 c^2 + m^2 c^4}$$

$$\text{Conservation of } \vec{p} \text{ implies } 2p = \frac{1}{c} \sqrt{E_1^2 + E_2^2 + 2E_1 E_2 \cos\theta}$$

$$\text{conservation of } E \text{ implies } 2\sqrt{p^2 c^2 + m^2 c^4} = E_1 + E_2$$

$$4p^2 c^2 = E_1^2 + E_2^2 + 2E_1 E_2 \cos\theta$$

$$4(p^2 c^2 + m^2 c^4) = E_1^2 + E_2^2 + 2E_1 E_2$$

$$4m^2 c^4 = 2E_1 E_2 (1 - \cos\theta) \Rightarrow E_1 E_2 = \frac{2m^2 c^4}{1 - \cos\theta}$$

$$\boxed{E_1 = \frac{2m^2 c^4}{E_2 (1 - \cos\theta)}}$$

E_1 will be minimized for a head-on collision

$$\cos\theta = -1 \Leftrightarrow \theta = \pi$$

In this case $E_1 = \frac{m^2 c^4}{E_2} = \frac{(mc^2)^2}{E_2}$

$$mc^2 \approx 0.5 \times 10^6 \text{ eV}$$

$$E_2 \approx 3 \text{ K} \cdot \frac{\frac{1}{40} \text{ eV}}{300 \text{ K}} = \frac{1}{4000} \text{ eV}$$

$$E_1 = \frac{0.25 \times 10^{12} (\text{eV})^2}{\frac{1}{4000} \text{ eV}} = \boxed{10^{15} \text{ eV}}$$

Question: Scattering

Consider the scattering of a spinless particle of mass m from a diatomic molecule. The incoming particle travels along the z -axis. Assume that the molecule is much heavier than the scattering particle and that there is no recoil. The two atoms in the molecule are aligned along the y -axis and localized at $y = b$ and $y = -b$. The potential the particle feels in the presence of the molecule can be modeled by the following potential:

$$V(\vec{x}) = \alpha \left(\delta(y - b) \delta(x) \delta(z) + \delta(y + b) \delta(x) \delta(z) \right)$$

- a) Calculate the scattering amplitude in the first Born approximation.
- b) Calculate the differential cross section from a) (Express the result in terms of the scattering angles).
- c) Calculate the total cross section. Do the integrals exactly. You might find the following integrals helpful:

$$\int_0^{2\pi} \alpha \cos(x \sin \alpha)^2 = \pi(1 + J_0(2x)), \quad \int_0^\pi d\alpha \sin \alpha J_0(x \sin \alpha) = \frac{2 \sin x}{x}$$

Solution:

a) The scattering amplitude in the first Born approximation is given by

$$f^{(1)}(\vec{k}', \vec{k}) = -\frac{2m(2\pi)^{3/2}}{4\pi\hbar^2} \int d^3y e^{-i(\vec{k}' - \vec{k})\vec{y}} V(\vec{y}) \quad (0.1)$$

We have

$$\vec{k} = \hat{e}_z k, \quad \vec{k}' = k(\hat{e}^z \cos \theta + \hat{e}_x \sin \theta \cos \phi + \hat{e}_y \sin \theta \sin \phi) \quad (0.2)$$

Plugging in the potential and evaluating the integral over y gives

$$f^{(1)}(\vec{k}', \vec{k}) = -\frac{2m(2\pi)^{3/2}\alpha}{4\pi\hbar^2} (e^{ikb \sin \theta \sin \phi} + e^{-ikb \sin \theta \sin \phi}) \quad (0.3)$$

$$= -\frac{m\alpha(2\pi)^{3/2}}{\hbar^2\pi} \cos(kb \sin \theta \sin \phi) \quad (0.4)$$

b) The differential cross section can be calculated from the scattering amplitude by

$$\begin{aligned} \frac{d\sigma}{d\Omega} &= |f(\theta, \phi)|^2 \\ &= \frac{\alpha^2 m^2 8\pi}{\hbar^4} |\cos(kb \sin \theta \sin \phi)|^2 \end{aligned} \quad (0.5)$$

c) The total cross section is given by

$$\begin{aligned} \sigma_{tot} &= \int d\Omega |f(\theta, \phi)|^2 \\ &= \frac{\alpha^2 m^2 8\pi}{\hbar^4} \int_0^\pi d\theta \sin \theta \int_0^{2\pi} d\phi |\cos(kb \cos \theta \cos \phi)|^2 \\ &= \frac{\alpha^2 m^2 8\pi^2}{\hbar^4} \int_0^\pi d\theta \sin \theta (1 + \pi J_0(2kb \sin \theta)) \\ &= \frac{\alpha^2 m^2 8\pi^2}{\hbar^4} \left(2 + \frac{\sin(2bk)}{bk}\right) \end{aligned} \quad (0.6)$$

Q4.

Fall 2013 Comprehensive Exam Questions

Question 4: Quantum Mechanics

Consider a system of two spin $1/2$ particles, labelled a and b , with respective spin operators \mathbf{S}_a and \mathbf{S}_b . We ignore all quantum numbers but those of spin. The particles are in the state $|\Psi\rangle$ of zero total angular momentum, which we consider normalized to $\langle\Psi|\Psi\rangle = 1$. Let \mathbf{n}_a and \mathbf{n}_b be two independent unit vectors. Compute the expectation value of the product of the spin operators projected onto the directions \mathbf{n}_a and \mathbf{n}_b respectively, namely,

$$\langle\Psi|(\mathbf{n}_a \cdot \mathbf{S}_a)(\mathbf{n}_b \cdot \mathbf{S}_b)|\Psi\rangle \quad (0.1)$$

Solution to Question 4:

Recall that the normalized spin 0 state $|\Psi\rangle$ of the two-spin system, in a basis where the spin operators S_a^z and S_b^z are diagonal with eigenstates $|\pm\rangle_a$ and $|\pm\rangle_b$, is given by,

$$|\Psi\rangle = \frac{1}{2}(|+\rangle_a \otimes |-\rangle_b - |-\rangle_a \otimes |+\rangle_b) \quad (0.2)$$

Also recall that the spin operators in this basis are given by $S_a^i = \hbar\sigma_a^i/2$ and $S_b^i = \hbar\sigma_b^i/2$, where σ_a^i and σ_b^i are the standard Pauli matrices acting in the Hilbert spaces for a and b .

The expectation value to be calculated is linear in \mathbf{n}_a and linear in \mathbf{n}_b . It is also invariant under rotations, since the state $|\Psi\rangle$ is invariant under rotations. Thus, the outcome must be proportional to the only possible rotation invariant bilinear in \mathbf{n}_a and \mathbf{n}_b ,

$$\langle\Psi|(\mathbf{n}_a \cdot \mathbf{S}_a)(\mathbf{n}_b \cdot \mathbf{S}_b)|\Psi\rangle = C \mathbf{n}_a \cdot \mathbf{n}_b \quad (0.3)$$

The coefficient C is independent of \mathbf{n}_a and \mathbf{n}_b . To evaluate it, we may choose any convenient assignment of \mathbf{n}_a and \mathbf{n}_b such that their inner product is non-zero. We take $\mathbf{n}_a = \mathbf{n}_b = (0, 0, 1)$. It is now straightforward to evaluate C by first computing, by inspection,

$$S_a^z S_b^z |\Psi\rangle = -\left(\frac{\hbar}{2}\right)^2 |\Psi\rangle \quad (0.4)$$

Since $|\Psi\rangle$ is normalized, we thus find,

$$C = -\left(\frac{\hbar}{2}\right)^2 \quad (0.5)$$

By the way, it is also easy to double check that the expectation value vanishes when $\mathbf{n}_a \cdot \mathbf{n}_b = 0$, by taking for example $\mathbf{n}_a = (1, 0, 0)$ and $\mathbf{n}_b = (0, 0, 1)$. We then find,

$$S_a^x S_b^z |\Psi\rangle = -\left(\frac{\hbar}{2}\right)^2 \frac{1}{2}(|-\rangle_a \otimes |-\rangle_b + |+\rangle_a \otimes |+\rangle_b) \quad (0.6)$$

whose inner product with $|\Psi\rangle$ vanishes term by term.

Problem 3

Consider a one-dimensional simple harmonic oscillator with the usual Hamiltonian,

$$H_0 = \frac{p^2}{2m} + \frac{1}{2} m \omega^2 x^2$$

Now, consider the effect of a perturbation,

$$V = \frac{1}{2} \epsilon m \omega^2 x^2 \quad \text{with} \quad \epsilon \ll 1$$

Find using perturbation theory, the new ground state ket $|0\rangle$ to order ϵ and the ground state energy shift to order ϵ^2 . Solve this problem exactly and compare with the results obtained using perturbation theory.

You may assume without proof that

$$\langle u_{n'} | x | u_n \rangle = \sqrt{\frac{\hbar}{2m\omega}} (\sqrt{n+1} \delta_{n', n+1} + \sqrt{n} \delta_{n', n-1})$$

where $|u_n\rangle$ is the n th eigenstate of H_0 .

Soln:-

The new ground state, $|0\rangle$ is given by

$$|0\rangle = |0^{(0)}\rangle + \sum_{\vec{k} \neq 0} |k^0\rangle \frac{V_{k0}}{E_0^{(0)} - E_k^{(0)}} + \dots$$

The ground state energy shift Δ_0 by

$$\Delta_0 = V_{00} + \sum_{k \neq 0} \frac{|V_{k0}|^2}{E_0^{(0)} - E_k^{(0)}} + \dots$$

Now

$$V_{00} = \frac{\epsilon m \omega^2}{2} \langle 0^{(0)} | x^2 | 0^{(0)} \rangle$$

$$= \sum_n \frac{\epsilon m \omega^2}{2} \langle 0^{(0)} | x | n^{(0)} \rangle \langle n^{(0)} | x | 0^{(0)} \rangle$$

$$= \sum_n \frac{\epsilon m \omega^2}{2} \left(\sqrt{\frac{\hbar}{2m\omega}} \delta_{n,1} \times \sqrt{\frac{\hbar}{2m\omega}} \delta_{1,n} \right)$$

$$= \frac{\epsilon \hbar \omega}{4}$$

For $k \neq 0$,

$$V_{k0} = \frac{\epsilon m \omega^2}{2} \sum_n \langle k^{(0)} | x | n^{(0)} \rangle \langle n^{(0)} | x | 0 \rangle$$

$$= \frac{\epsilon m \omega^2}{2} \times \left(\frac{\hbar}{2m\omega} \right) \times \sum_n \sqrt{2} \cdot \delta_{k,2} \times \delta_{1,n}$$

$$= \frac{\epsilon \hbar \omega}{2\sqrt{2}} \delta_{k,2}$$

Thus, $|0\rangle = |0^{(0)}\rangle - \frac{\epsilon}{4\sqrt{2}} |2^{(0)}\rangle + O(\epsilon^2)$

$$\Delta_0 = E_0 - E_0^{(0)} = \hbar\omega \left[\frac{\epsilon}{4} - \frac{\epsilon^2}{16} + O(\epsilon^3) \right]$$

The exact solution for $H_0 + V$ is obtained by substitution $\omega \rightarrow \omega\sqrt{1+\epsilon}$ in the solution for H_0 .

Thus ground state energy

$$= \frac{\hbar\omega\sqrt{1+\epsilon}}{2} = \frac{\hbar\omega}{2} \left[1 + \frac{\epsilon}{2} - \frac{\epsilon^2}{8} + \dots \right]$$

The ground state wave function of H_0 is

$$\langle x|0 \rangle = \frac{1}{\pi^{1/4}} \frac{1}{\sqrt{x_0}} e^{-x^2/2x_0^2} \quad \text{where} \quad x_0 = \sqrt{\frac{\hbar}{m\omega}}$$

Hence the G.S wavefn of $H_0 + V$ is

obtained by $x_0 \rightarrow \frac{x_0}{(1+\epsilon)^{1/4}}$.

Hence $\langle x|0 \rangle \rightarrow \frac{1}{\pi^{1/4}\sqrt{x_0}} (1+\epsilon)^{1/8} \exp \left[- \left(\frac{x^2}{2x_0^2} \right) (1+\epsilon)^{1/2} \right]$

$$\approx \frac{1}{\pi^{1/4}} \frac{1}{\sqrt{x_0}} e^{-\frac{x^2}{2x_0^2}} + \frac{\epsilon}{\pi^{1/4}\sqrt{x_0}} e^{-\frac{x^2}{2x_0^2}} \left[\frac{1}{8} - \frac{1}{4} \frac{x^2}{x_0^2} \right]$$

$$\text{Now, } \langle x | 2^{(0)} \rangle = \frac{1}{\sqrt{2}} (a^+)^2 |0\rangle, \text{ with}$$

$$a^+ = \sqrt{\frac{m\omega}{2\hbar}} \left(x - \frac{ip}{m\omega} \right) = \frac{1}{\sqrt{2}x_0} \left(x - x_0^2 \frac{d}{dx} \right)$$

$$\text{Thus, } \langle x | 2^{(0)} \rangle = \frac{1}{\sqrt{2}} \left(\frac{1}{\sqrt{2}x_0} \right)^2 \left(x - x_0^2 \frac{d}{dx} \right)^2 \langle x | 0 \rangle$$

$$= \frac{1}{2\sqrt{2}} \cdot \frac{1}{\pi^{1/4}} \frac{1}{\sqrt{x_0}} e^{-x^2/2x_0^2} \left[-2 + 4 \left(\frac{x}{x_0} \right)^2 \right]$$

$$\therefore \langle x | 0 \rangle = \langle x | 0^{(0)} \rangle - \frac{\epsilon}{4\sqrt{2}} \langle x | 2^{(0)} \rangle$$

QUESTION: Perturbation theory time (in)dependent

The Hamiltonian of a particle of mass m in a one dimensional infinite well is given by

$$H_0 = \frac{p^2}{2m} + V(x), \quad V(x) = \begin{cases} 0 & 0 < x < L \\ \infty & \text{otherwise} \end{cases}$$

A time dependent perturbation is added

$$H_{\text{total}} = H_0 + H', \quad H' = \lambda \delta(x - L/2) f(t)$$

Where λ is constant and $f(t)$ is a time dependent function.

a) Calculate the matrix elements of H' with the eigenstates of the unperturbed Hamiltonian.

b) In this part of the problem we consider a time independent perturbation $f(t) = 1$.

First, calculate the first nonzero correction of the ground state energy. Second, at which order (if any) will the first excited state receive a nonzero correction? Please back up your answer with an argument/calculation.

c) Now take the following time dependent function

$$f(t) = \int_{-\infty}^{\infty} d\omega \rho(\omega) (e^{i\omega t} + e^{-i\omega t})$$

with

$$\rho(\omega) = \sqrt{\frac{\alpha}{\pi}} e^{-\alpha\omega^2}$$

If the system is in its ground state at time $t = 0$, what are the possible transitions into excited states that the system can make at time $t > 0$? Use first order time-dependent perturbation theory.

d) Find the transition rate into these excited states using Fermi's golden rule.

Solution:

The free particle in a box has eigenfunctions

$$\psi_n(x) = \sqrt{\frac{2}{L}} \sin \frac{n\pi x}{L}, \quad n \in N \quad (0.7)$$

With eigenvalues of H_0 given by

$$E_n = \frac{\hbar^2}{2m} \frac{n^2 \pi^2}{L^2} \quad (0.8)$$

a) The matrix elements are

$$\begin{aligned} \langle n | H' | m \rangle &= \frac{2}{L} \lambda f(t) \int_0^L dx \sin \frac{n\pi x}{L} \sin \frac{m\pi x}{L} \delta(x - L/2) \\ &= \frac{2\lambda f(t)}{L} \sin \frac{n\pi}{2} \sin \frac{m\pi}{2} \\ &= \begin{cases} \frac{2\lambda f(t)}{L} & n, m \text{ odd} \\ 0 & \text{otherwise} \end{cases} \end{aligned} \quad (0.9)$$

b) For the ground state one has $n = 1$ which is odd, hence the first nonzero correction occurs at first order in perturbation theory and is given by

$$\Delta E_1^{(1)} = \langle 1 | H' | 1 \rangle = \frac{2\lambda}{L} \quad (0.10)$$

For the first excited state one has $n = 2$, in the formal power series of time independent non degenerate perturbation theory the m -th order contribution to the energy E_2 is given by

$$\Delta E_2^{(m)} = \langle 2_{(0)} | H' | 2^{(m-1)} \rangle \quad (0.11)$$

Where $| 2^{(m-1)} \rangle$ is the wave function correction to order $m - 1$. No matter which form $| 2^{(m-1)} \rangle$ takes, the inner product always vanishes since $\psi_2(x)$ has a zero at $x = L/2$. Hence the first excited state is not corrected to all orders in perturbation theory.

c) The generate time dependent wave function van be expanded as

$$| \psi(t) \rangle = \sum_n c_n(t) e^{-i \frac{1}{\hbar} E_n t} | n \rangle \quad (0.12)$$

With

$$c_n(t) = c_n(0) - \frac{i}{\hbar} \sum_{m \neq n} \int_0^t dt' \langle n | H' | m \rangle e^{-i \frac{(E_m - E_n)t'}{\hbar}} c_m(t') \quad (0.13)$$

In first order time dependent perturbation theory $c_m(t')$ gets replaced by $c_m(0)$. Since the initial condition sets $c_m(0) = 0$ for $m \neq 1$ the only nonzero matrix element occurs for odd $n = 1, 3, 5, \dots$, because of (0.23). Hence the only transitions can occur to states with odd n .

d) For an interaction of the form

$$H' = V(e^{i\omega t} + e^{-i\omega t}) \quad (0.14)$$

Fermi's Golden rule gives the transition rate

$$\Gamma(\omega)_{1 \rightarrow 2k+1} = \frac{2\pi}{\hbar} |\langle 1|V|2k+1 \rangle|^2 \left(\delta(E_{2k+1} - E_1 - \hbar\omega) + \delta(E_{2k+1} - E_1 + \hbar\omega) \right) \quad (0.15)$$

With

$$V = \lambda \delta(x - L/2) \quad (0.16)$$

One finds from part a)

$$|\langle 1|V|2k+1 \rangle|^2 = \frac{4\lambda^2}{L^2} \quad (0.17)$$

The total rate then becomes

$$\begin{aligned} \Gamma_{tot} &= \int_{-\infty}^{\infty} d\omega \rho(\omega) \Gamma(\omega)_{1 \rightarrow 2k+1} \\ &= \frac{2\pi}{\hbar} \left\{ \rho\left(\frac{1}{\hbar}(E_{2k+1} - E_1)\right) + \rho\left(-\frac{1}{\hbar}(E_{2k+1} - E_1)\right) \right\} |\langle 0|V|2k+1 \rangle|^2 \\ &= \frac{2\pi}{\hbar} \frac{4\lambda^2}{L^2} \sqrt{\frac{\alpha}{\pi}} 2e^{-\frac{\alpha}{\hbar^2}(E_{2k+1} - E_1)^2} \end{aligned} \quad (0.18)$$

Quantum Mechanics Q1 Jaynes-Cummings Model

A two-level system with bare eigenstates $|\uparrow\rangle$ and $|\downarrow\rangle$ ("a qubit") has an energy splitting ω_q ($\hbar \equiv 1$). Consider the coupling between the qubit and a harmonic oscillator (bare eigenstates $|n\rangle$ for $n \in \{0, 1, 2, \dots\}$) with energy splitting ω_0 , which is described by the Hamiltonian

$$H = H_0 + H_{\text{int}}$$

$$H_0 = \omega_0 (\hat{a}^\dagger \hat{a} + \frac{1}{2}) + \frac{1}{2} \omega_q \hat{\sigma}_z$$

$$H_{\text{int}} = \frac{\Omega}{2} (\hat{a}^\dagger \hat{\sigma}_- + \hat{a} \hat{\sigma}_+)$$

Where Ω is the coupling strength, \hat{a}^\dagger and \hat{a} are the harmonic oscillator raising and lowering operators, and the spin- $\frac{1}{2}$ operators are defined as $\hat{\sigma}_z \equiv |\uparrow\rangle\langle\uparrow| - |\downarrow\rangle\langle\downarrow|$, $\hat{\sigma}_+ \equiv |\uparrow\rangle\langle\downarrow|$, $\hat{\sigma}_- \equiv |\downarrow\rangle\langle\uparrow|$.

a.) Consider the "dispersive regime," defined by $\omega_0 - \omega_q \gg \frac{\Omega}{2}$.

Find the energies and eigenstates of H . You may treat H_{int} as a perturbation and use 1st-nonzero-order terms. Consider coupling between nearly-degenerate states only. Let $\Delta \equiv \omega_0 - \omega_q > 0$. Don't bother to normalize the eigenstates.

b.) Consider the "resonant regime," defined by $\omega_q \approx \omega_0 \equiv \omega$.

Find the energies and eigenstates of H . You may treat H_{int} as a perturbation and use 1st-order degenerate perturbation theory. Hint: consider the state $\frac{1}{\sqrt{2}}(|\downarrow, n\rangle + e^{i\varphi} |\uparrow, n-1\rangle)$. Consider coupling between degenerate states only.

a.) Solution #1: plain-old perturbation theory

$$E_{\downarrow, n}^0 = n\omega_0 + \frac{\omega_0}{2} - \frac{\omega_0}{2} = n\omega_0 + \frac{\Delta}{2}$$

$$E_{\uparrow, n-1}^0 = n\omega_0 - \frac{\Delta}{2}$$

Apply Merzbacher 8.37 or Sakurai 5.1.11

$$E_{+, n} = \frac{1}{2} \left[2n\omega_0 + \sqrt{(-\Delta)^2 + 4 \left| \frac{\Omega}{2} \sqrt{n} \right|^2} \right]$$

$$= n\omega_0 + \frac{1}{2} \Delta \sqrt{1 + \frac{4 \left| \frac{\Omega}{2} \sqrt{n} \right|^2}{\Delta^2}}$$

$$\approx 1 + \frac{n\Omega^2}{2\Delta^2}$$

$$\approx \left[n\omega_0 + \frac{1}{2} \Delta + \frac{n\Omega^2}{4\Delta} \right]$$

$$E_{-, n} = \frac{1}{2} \left[2n\omega_0 - \frac{\Delta}{2} \sqrt{1 + \frac{\Omega^2 n}{\Delta^2}} \right]$$

$$\approx \left[n\omega_0 - \frac{1}{2} \Delta - \frac{n\Omega^2}{4\Delta} \right]$$

Un-normalized eigenstates:

$$|t, n\rangle = |\downarrow, n\rangle + \frac{\langle \uparrow, n-1 | H_{int} | \downarrow, n \rangle}{E_{\downarrow, n}^0 - E_{\uparrow, n-1}^0} |\uparrow, n-1\rangle$$

$$= \left[|\downarrow, n\rangle + \frac{g\sqrt{n}}{\Delta} |\uparrow, n-1\rangle \right]$$

$$|-, n\rangle = |\uparrow, n-1\rangle - \frac{g\sqrt{n}}{\Delta} |\downarrow, n\rangle$$

Solution to QM1 Jaynes-Cummings Model

- Solution #1: eigenstates by inspection
b2)

Since we are instructed to consider coupling between degenerate states only (1st order), we have a series of two-level systems described by $|\downarrow, n\rangle$ and $|\uparrow, n-1\rangle$.

These have the same energy under H_0 ($n\omega$), so we will subtract this constant from the Hamiltonian. We note that H_{int} is completely off-diagonal in this basis

$$\langle \downarrow, n | H_{int} | \downarrow, n \rangle = \langle \uparrow, n-1 | H_{int} | \uparrow, n-1 \rangle = 0, \quad \langle \downarrow, n | H_{int} | \uparrow, n-1 \rangle = \frac{\Omega}{2}\sqrt{n}$$

So we are asked to diagonalize a 2×2 matrix that is proportional to $\hat{\sigma}_x$:

$$H_n = \frac{\Omega}{2}\sqrt{n} \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}$$

The eigenstates are therefore given by

b2)

$$\begin{aligned} |+, n\rangle &\equiv (|\downarrow, n\rangle + |\uparrow, n-1\rangle) \frac{1}{\sqrt{2}} \\ |-, n\rangle &\equiv (|\downarrow, n\rangle - |\uparrow, n-1\rangle) \frac{1}{\sqrt{2}} \end{aligned}$$

The energies can be found as follows

$$H_n |+, n\rangle = E_{+n} |+, n\rangle = \frac{\Omega}{2}\sqrt{n} |+, n\rangle$$

$$H_n |-, n\rangle = E_{-n} |-, n\rangle = -\frac{\Omega}{2}\sqrt{n} |-, n\rangle$$

b2) so the energies are given by $n\omega \pm \frac{\Omega}{2}\sqrt{n}$

Solution #2 : degenerate perturbation theory

b.)

Using, e.g., Merzbacher 8.37 we can write

$$E_{\pm} = \frac{1}{2} \left[n\omega + n\omega \pm \sqrt{\underbrace{(n\omega - n\omega)^2}_0 + 4 |\langle \downarrow, n | H_{int} | \uparrow, n-1 \rangle|^2} \right]$$

since $\langle \downarrow, n | H_{int} | \uparrow, n-1 \rangle = \frac{\sqrt{2}}{2} \sqrt{n}$, we have

b.)

$$\begin{cases} E_+ = n\omega + \frac{\sqrt{2}}{2} \sqrt{n} \\ E_- = n\omega - \frac{\sqrt{2}}{2} \sqrt{n} \end{cases}$$

Another way to get this is to notice the hint about the states $|\pm\rangle \equiv \frac{1}{\sqrt{2}} (|n, \downarrow\rangle \pm |n-1, \uparrow\rangle)$. Do these states diagonalize H_{int} ? Well, $H_{int} |t, n\rangle = \frac{\sqrt{2}}{2} \sqrt{n} |t, n\rangle$, so yes, they do. We can now use them to apply Sakurai 5.2.11:

$$\Delta_+ = \langle +, n | H_{int} | +, n \rangle = \frac{\sqrt{2}}{2} \sqrt{n}$$

$$\Delta_- = \langle -, n | H_{int} | -, n \rangle = -\frac{\sqrt{2}}{2} \sqrt{n}$$

b.) therefore,

$$\begin{aligned} E_+ &= \langle +, n | H_0 | +, n \rangle + \frac{\sqrt{2}}{2} \sqrt{n} = n\omega + \frac{\sqrt{2}}{2} \sqrt{n} \\ E_- &= \langle -, n | H_0 | -, n \rangle - \frac{\sqrt{2}}{2} \sqrt{n} = n\omega - \frac{\sqrt{2}}{2} \sqrt{n} \end{aligned}$$

Since we know that $|\pm, n\rangle$ diagonalize H_{int} , we can check to see if they're also eigenstates of H_0 :

$H_0 |\pm, n\rangle = n\omega$, so we have

$$\begin{cases} |+, n\rangle = \frac{1}{\sqrt{2}} (|n, \downarrow\rangle + |n-1, \uparrow\rangle) \\ |-, n\rangle = \frac{1}{\sqrt{2}} (|n, \downarrow\rangle - |n-1, \uparrow\rangle) \end{cases}$$

Question 8: Statistical Mechanics

We propose to evaluate the *Richardson effect*, namely the electric current density of electrons which is produced by heating up a metal in the presence of an external electric potential. The potential energy of an electron just outside the metal is denoted $W > 0$.

The potential energy for electrons inside the metal is taken to be 0. The electrons are considered otherwise non-interacting, and filled up to chemical potential μ with $\mu < W$. Since we consider the problem to be at sufficiently low temperature, μ may be identified with the Fermi energy.

- State the condition on the momentum of an electron that can escape from the metal to the outside as a function of W and μ .
- Derive a general expression for the current density I of electrons leaving the metal.
- Obtain an approximation of your result in (b) valid for sufficiently low temperatures.

Solution to Question 8

(a) We take the edge of the metal where the electrons are being emitted to be orthogonal to the z -direction. The condition for an electron to be able to escape the metal to the outside is that its kinetic energy in the z -direction can overcome the potential energy outside the metal, so that we then must have,

$$\frac{p_z^2}{2m} > W \quad (0.19)$$

where m is the electron mass, and p_z is the electron momentum in the z -direction.

(b) The density of electrons inside the metal in an infinitesimal phase space volume $dV d^3p$ (where dV is the spacial volume element) is given by,

$$2 \frac{dV d^3p}{(2\pi\hbar)^3} \frac{1}{e^{\beta(p^2/2m - \mu)} + 1} \quad (0.20)$$

The factor of 2 arises from the two spin states of the electron, and we have $\mathbf{p}^2 = p_x^2 + p_y^2 + p_z^2$. The electric current density is then given by the thermal expectation value of the observable,

$$\frac{e p_z}{m} \quad (0.21)$$

per unit volume, restricted to the range $p_z > \sqrt{2mW}$. Thus the current density $I = I_z$ is given by the following integral,

$$I_z = 2 \frac{e}{m} \frac{1}{(2\pi\hbar)^3} \int_{\sqrt{2mW}}^{\infty} dp_z \int_{-\infty}^{\infty} dp_x \int_{-\infty}^{\infty} dp_y \frac{p_z}{e^{\beta(p^2/2m - \mu)} + 1} \quad (0.22)$$

Changing variables to the following dimensionless combinations s, t defined by,

$$s = \beta \left(\frac{p_z^2}{2m} - W \right) \quad t = \beta \frac{p_x^2 + p_y^2}{2m} \quad (0.23)$$

The integral for I_z reduces to,

$$I_z = \frac{e m}{2\pi^2 \hbar^3} (kT)^2 \int_0^\infty ds \int_0^\infty dt \frac{1}{e^{s+t+\beta(W-\mu)} + 1} \quad (0.24)$$

(c) For sufficiently low temperatures, namely $T \ll W - \mu$, we may drop the 1 in the denominator, and carry out the integrals over s and t explicitly. We are then left with the following approximate formula,

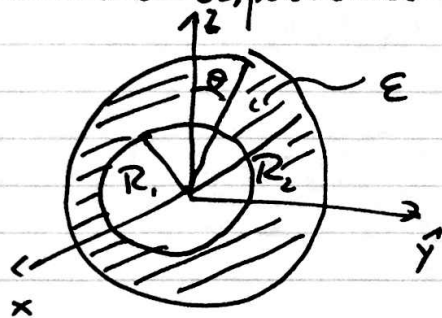
$$I_z = \frac{e m}{2\pi^2 \hbar^3} (kT)^2 \exp \left\{ -\frac{W - \mu}{kT} \right\} \quad (0.25)$$

1. The gap between two spherical conducting shells - ^①
 see the figure below - is filled with a spatially inhomogeneous dielectric so that the dielectric constant depends on the polar angle θ as:

$$\epsilon(\theta) = \epsilon_1 + \epsilon_2 \cos^4 \theta$$

a) When charged so that the inner and outer spheres have charges $+Q$ and $-Q$ respectively, show that the internal electric field is purely radial $\vec{E} = \hat{r} E_r$ and is independent of the angle ϕ, θ .

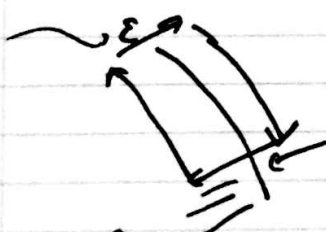
b) Calculate the capacitance C of the system.



Ans.

a) no ϕ -dependence by azimuthal symmetry: $E_\phi = 0$ and $\nabla \times \vec{E} = 0$
 Consider the line-integral shown below.

line
 integral



$$\nabla \times \vec{E} = 0 \Rightarrow \oint \vec{E} \cdot d\vec{l} = 0 \text{ and since}$$

\vec{E} vanishes in the conductor, E_θ must vanish in the dielectric.

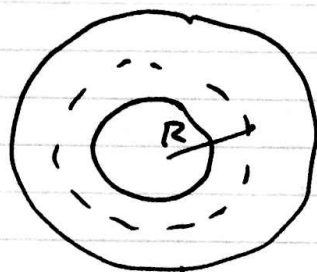
dielectric

$\Rightarrow E_\phi = 0$ the same argument can be extended throughout the gap.

$$\Rightarrow \vec{E} = \hat{r} E_r(r)$$

any other dependence $\Rightarrow \nabla \times \vec{E} \neq 0$.

b) Using Gauss's law inside.



$$R_1 < R < R_2$$

$$\oint \vec{D} \cdot d\vec{A} = 4\pi Q \quad \text{or}$$

$$2\pi R^2 E \int_{-1}^{+1} d(\cos\theta) \{ \epsilon_1 + \epsilon_2 \cos\theta \} = 4\pi Q$$

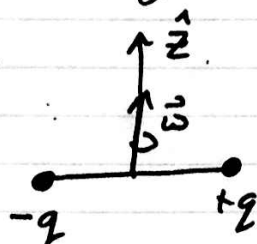
$$4\pi R^2 E \frac{(\epsilon_2 + \epsilon_1)}{5} = 4\pi Q \Rightarrow E = \frac{5Q}{R^2 (\epsilon_2 + \epsilon_1)}$$

$$\Delta V = - \int_{R_1}^{R_2} d\vec{r} \cdot \vec{E} = \frac{5Q}{(\epsilon_2 + \epsilon_1)} \left(\frac{1}{R_1} - \frac{1}{R_2} \right)$$

$$C = \frac{Q}{\Delta V} = \frac{(\epsilon_2 + \epsilon_1)}{5} \left[\frac{R_2 R_1}{R_2 - R_1} \right]$$

$$\int_{-1}^{+1} x^4 dx = \frac{1}{5} x^5 \Big|_{-1}^{+1} = 2/5$$

3. A small stick of length l is hard rotating in the xy plane. (5)
 This stick has charges $\pm q$ embedded on either end - see figure below.



$\vec{\omega} = \hat{z} \omega_0$ initial angular velocity.
 (moment of inertia I .)

You observe that its angular velocity is slowly decreasing:
 $\dot{\omega} < 0$ and $|\dot{\omega}| \ll \omega$. Predict $\omega(t)$.

Ans. Slowing due to radiation.

Rotating dipole $\vec{p} = (\hat{x} + i\hat{y}) q l e^{-i\omega t}$ (5)

Power radiated $\frac{dP}{d\Omega} = \frac{c k^4}{8\pi} |(\hat{r} \times \hat{p}) \times \hat{r}|^2$; $k = \omega/c$

$$(\hat{r} \times \hat{p}) \times \hat{r} = \vec{p} - (\hat{r} \cdot \vec{p}) \hat{r}$$
(5)

$$\hat{r} \cdot \vec{p} = z q l [\sin\theta \cos\phi + i \sin\theta \sin\phi] = z q l \sin\theta e^{i\phi}$$

$$|(\hat{r} \times \hat{p}) \times \hat{r}|^2 = 4 q^2 l^2 [\hat{x} + i\hat{y} - \sin\theta e^{i\phi} \hat{r}] \cdot [\hat{x} - i\hat{y} - \sin\theta e^{-i\phi} \hat{r}]$$

$$= 2(2 - \sin^2\theta) \cdot 4 q^2 l^2 \text{ After some algebra...}$$

$$\frac{dP}{d\Omega} = \frac{\omega^4}{2c^3\pi} [1 + \cos^2\theta]; \text{ Integrate over the unit sphere}$$

to get total power radiated:

$$P = \oint \frac{dP}{d\Omega} d\Omega = \frac{2\pi\omega^4}{2c^3\pi} \left[x + \frac{x^3}{3} \right]_{-1}^1 = \frac{8}{3} \frac{\omega^4}{c^3} (q l)^2$$
(5)

⑥

$$\text{Now } \frac{d}{dt} \left(\frac{1}{2} I \dot{\omega}^2 \right) = -P ;$$

$$\Rightarrow I \dot{\omega} = -P = -A \omega^4$$

$$I \dot{\omega} = -A \omega^3 \quad \text{or} \quad \int \frac{d\omega}{\omega^3} = - \frac{A}{I} \int_0^t dt$$

$\omega_0 \leftarrow \text{initial}$

$$-\frac{1}{2} \frac{1}{\omega^2} \bigg|_{\omega_0}^{\omega} = - \frac{A}{I} t$$

$$\frac{1}{\omega^2} - \frac{1}{\omega_0^2} = \frac{2A}{I} t \Rightarrow \frac{1}{\omega^2} = \frac{I}{\omega_0^2 I} + \frac{2At\omega_0^2}{I\omega_0^2}$$

$$\omega(t) = \sqrt{\frac{\omega_0^2 I}{I + 2At\omega_0^2}} ; \quad A = \frac{8}{3c^3} (ql)^2$$

+5

Q12

Assume the existence of magnetic charge related to the magnetic field by the local reaction

$$\vec{\nabla} \cdot \vec{B} = \mu_0 \rho_m.$$

(a). Using the Gauss's theorem, obtain the magnetic field \vec{B} of a point magnetic charge at the origin.

(b). In the absence of the magnetic charge, the curl of the electric field is given by the

Faraday's law, $\vec{\nabla} \times \vec{E} = -\frac{\partial \vec{B}}{\partial t}$. Show that this law is incompatible with the magnetic charge density that is a function of time.

(c). Assuming that magnetic charge is conserved, derive the local relation between the magnetic charge current density \vec{J}_m and the magnetic density ρ_m .

(d). Modify Faraday's law as given in part (b) to obtain a law consistent with the presence of the magnetic charge density that is a function of position and time.

Solution

$$(a) \int_V \vec{\nabla} \cdot \vec{B} dV = \oint_S \vec{B} \cdot d\vec{S} = 4\pi r^2 B(r) = \mu_0 q_m$$

$$\vec{B}(r) = \frac{\mu_0 q_m}{4\pi r^2} \hat{r}$$

$$(b) \frac{\partial}{\partial t} \vec{\nabla} \cdot \vec{B} = \vec{\nabla} \cdot \frac{\partial \vec{B}}{\partial t} = -\vec{\nabla} \cdot (\vec{\nabla} \times \vec{E}) = 0$$

$$\text{On the other hand, } \frac{\partial}{\partial t} \vec{\nabla} \cdot \vec{B} = \mu_0 \frac{\partial \rho_m}{\partial t}$$

Thus the Faraday's law is incompatible with the magnetic charge density that is a function of time.

$$(c) \frac{\partial}{\partial t} \int_V \rho_m dV = -\oint_S \vec{J}_m \cdot d\vec{S} = -\oint_S \vec{\nabla} \cdot \vec{J}_m dV$$

$$\frac{\partial \rho_m}{\partial t} + \vec{\nabla} \cdot \vec{J}_m = 0$$

This is the continuity equation for magnetic charge.

$$(d) \text{ If we modify Faraday's law, } \vec{\nabla} \times \vec{E} = -\mu_0 \vec{J}_m - \frac{\partial \vec{B}}{\partial t}$$

$$\text{and } -\mu_0 \vec{\nabla} \cdot \vec{J}_m - \frac{\partial \vec{\nabla} \cdot \vec{B}}{\partial t} = -\mu_0 \left(\vec{\nabla} \cdot \vec{J}_m + \frac{\partial \rho_m}{\partial t} \right) = 0$$

$$\text{Hence } \frac{\partial}{\partial t} \vec{\nabla} \cdot \vec{B} = -\mu_0 \vec{\nabla} \cdot \vec{J}_m = \mu_0 \frac{\partial \rho_m}{\partial t} \text{ which is consistent with the second equation in (b).}$$

Q13

An un-polarized plane electromagnetic wave is scattered by a free electron. Derive the differential cross-section for scattering in the non-relativistic limit (Thompson scattering).

Solution

Consider an incident plane wave

$$\vec{E}_i = E_0 e^{-i(\omega t - \vec{k} \cdot \vec{x})} \hat{e}_0$$

The force on the free electron is

$$\vec{F} = -e\vec{E}_i \sim -eE_0 e^{-i\omega t} \hat{e}_0 = m\ddot{\vec{x}} = -m\omega^2 \vec{x} \quad \vec{x} = \frac{eE_0 \hat{e}_0}{m\omega^2} e^{-i\omega t}$$

The induced dipole moment is $\vec{p} = -e\vec{x} = -\frac{e^2 E_0 \hat{e}_0}{m\omega^2} e^{-i\omega t}$

The scattered electric field is $\vec{E}_s = \frac{k^2}{4\pi\epsilon_0} \frac{e^{ikr}}{r} [(\hat{n} \times \vec{p}) \times \hat{n}]$

The differential cross section is $\frac{d\sigma}{d\Omega} = \frac{k^4}{(4\pi\epsilon_0)^2} \left(\frac{q^2}{m\omega^2} \right)^2 |\hat{n} \times \hat{e}_0|^2$ where $\omega = ck$.

$$|\hat{n} \times \hat{e}_0|^2 = \begin{vmatrix} \hat{x} & \hat{y} & \hat{z} \\ n_x & n_y & n_z \\ \sin\phi & \cos\phi & 0 \end{vmatrix}^2 = \left| -\hat{x}n_z \cos\phi + \hat{y}n_z \sin\phi + \hat{z}(n_x \cos\phi - n_y \sin\phi) \right|^2$$

$$= n_z^2 (\cos^2\phi + \sin^2\phi) + n_x^2 \cos^2\phi + n_y^2 \sin^2\phi - 2n_x n_y \sin\phi \cos\phi$$

Using $\langle \cos^2\phi \rangle = \langle \sin^2\phi \rangle = 1/2$ and $\langle \cos\phi \sin\phi \rangle = 0$, we have

$$|\hat{n} \times \hat{e}_0|^2 = n_z^2 + \frac{1}{2}(n_x^2 + n_y^2) = \cos^2\theta + \frac{1}{2}\sin^2\theta = \frac{1}{2}(1 + \cos^2\theta)$$

$$\text{Thus we obtain} \quad \frac{d\sigma}{d\Omega} = \frac{1}{(4\pi\epsilon_0)^2} \left(\frac{e^2}{mc^2} \right)^2 \frac{1 + \cos^2\theta}{2} = \frac{1 + \cos^2\theta}{2} r_e^2$$

Where $r_e \equiv \frac{e^2}{4\pi\epsilon_0 mc^2}$ is the classical electron radius.

Q14
In cgs units

$$\textcircled{1} \nabla \cdot \vec{D} = 4\pi\rho$$

$$\vec{D} = \epsilon \vec{E}$$

$$\textcircled{2} \nabla \cdot \vec{B} = 0$$

$$\vec{B} = \mu \vec{H}$$

$$\textcircled{3} \nabla \times \vec{H} = \frac{4\pi}{c} \vec{J} + \frac{1}{c} \frac{\partial \vec{D}}{\partial t}$$

$$\textcircled{4} \nabla \times \vec{E} = -\frac{1}{c} \frac{\partial \vec{B}}{\partial t}$$

assume $\vec{E} = \vec{E}_0 e^{i(\vec{k} \cdot \vec{r} - \omega t)}$

Then $\textcircled{4}$ gives $\vec{k} \times \vec{E} = \frac{\omega}{c} \vec{B}$

$$\Rightarrow \vec{B} = \hat{k} \times \vec{E}$$

Normally $\vec{J} = \nabla \times \vec{E}$. Here we have a surface current

$$\vec{K} = \nabla_{2D} \times \vec{E}$$

where $\sigma_{2D} = \frac{1}{Z_0} = \frac{1}{377 \Omega} \Big|_{\text{MKS}} = \frac{c}{4\pi} \Big|_{\text{cgs}}$

The boundary conditions come from $\textcircled{4}$ and $\textcircled{3}$.

From $\textcircled{4}$ $\oint \vec{E} \cdot d\vec{l} = -\frac{1}{c} \frac{\partial}{\partial t} \int \vec{B} \cdot d\vec{a} = 0$ \vec{E} is continuous

$$\vec{E}_i + \vec{E}_r = \vec{E}_t \quad \text{as drawn} \Rightarrow \boxed{E_i - E_r = E_t}$$

From $\textcircled{3}$ $\oint \vec{H} \cdot d\vec{l} = \oint \vec{B} \cdot d\vec{l} = \frac{4\pi}{c} \int \vec{J} \cdot d\vec{a} + \frac{1}{c} \frac{\partial}{\partial t} \int \vec{D} \cdot d\vec{a}$

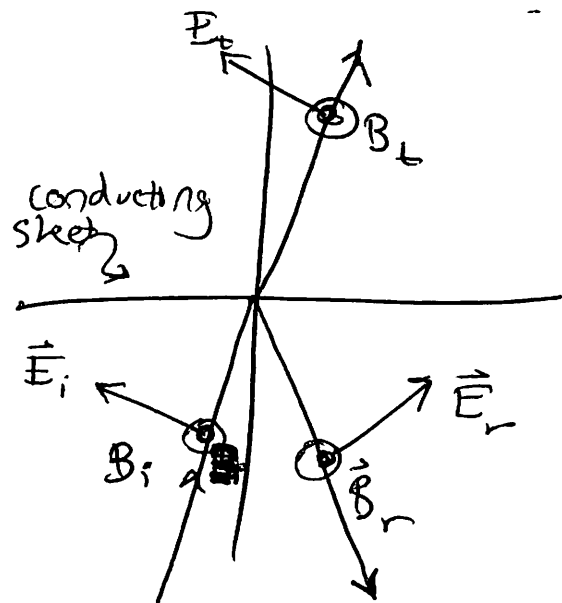
$$B_i + B_r - B_t = \frac{4\pi}{c} K = \frac{4\pi}{c} \sigma_{2D} E_t$$

loop has 0 area

using $|\vec{B}| = |\vec{E}|$

$$\boxed{E_i + E_r - E_t = S E_t}$$

where $S \equiv \frac{4\pi}{c} \sigma_{2D}$
 $= Z_0 \sigma_{2D}$



Drawn with finite angle of incidence for clarity

With the boxed equations

Eliminate E_r

$$E_i + (E_i - E_t) - E_t = S E_t$$

$$2(E_i - E_t) = S E_t$$

$$E_i = \left(\frac{S}{2} + 1\right) E_t$$

$$\frac{E_t}{E_i} = \frac{1}{\frac{S}{2} + 1}$$

Eliminate E_t

$$E_i + E_r - (E_i - E_r) = S(E_i - E_r)$$

$$2E_r = S(E_i - E_r)$$

$$E_r(1 + \frac{S}{2}) = \frac{S}{2} E_i$$

$$\frac{E_r}{E_i} = \frac{\frac{S}{2}}{\frac{S}{2} + 1}$$

$$T = \left(\frac{E_t}{E_i}\right)^2 = \frac{1}{(1 + \frac{S}{2})^2}$$

$$R = \left(\frac{E_r}{E_i}\right)^2 = \left(\frac{\frac{S}{2}}{1 + \frac{S}{2}}\right)^2$$

$$T + R + A = 1 \quad \therefore A = \frac{S}{(1 + \frac{S}{2})^2}$$

For a sheet impedance of 377Ω , $S = 1$ and

$$\boxed{T = \frac{4}{9} \quad R = \frac{1}{9} \quad A = \frac{4}{9}}$$

To maximize A

$$\frac{dA}{dS} = \frac{1}{(1 + \frac{S}{2})^2} - \frac{2S \cdot \frac{1}{2}}{(1 + \frac{S}{2})^3} = 0 \quad 1 = \frac{S}{1 + \frac{S}{2}}$$

$$1 + \frac{S}{2} = S \quad \frac{S}{2} = 1 \quad S = 2$$

So we need $\sigma_{20} = \frac{2}{Z_0}$, or a sheet resistivity of $\frac{377}{2} \Omega = \boxed{189 \Omega}$, this gives $\boxed{A = \frac{1}{2}}$

Half of the electromagnetic wave's energy is carried in the \vec{B} field, while the thin sheet only couples dissipatively to the \vec{E} field. Thus even with the best impedance match it can only achieve 50% absorption.